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Transition to turbulence in Hunt's flow in a moderate magnetic field

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Abstract – Pressure-driven magnetohydrodynamic duct flow in a transverse uniform magnetic field is studied by direct numerical simulation. The electric boundary conditions correspond to Hunt's flow with perfectly insulating walls parallel to the magnetic field (side walls) and perfectly conducting walls perpendicular to the magnetic field (Hartmann walls). The velocity distribution exhibits strong jets at the side walls, which are susceptible to instability even at low Reynolds numbers Re. We explore the onset of time-dependent flow and transition to states with evolved turbulence for a moderate Hartmann number Ha = 100. At low Re time-dependence appears in the form of elongated Ting-Walker vortices at the side walls of the duct, which, upon increasing Re, develop into more complex structures with higher energy and then the side-wall jets partially detach from the walls. At high values of Re jet detachments disappear and the flow consists of two turbulent jets and nearly laminar core. It is also demonstrated that, there is a range of Re, where Hunt's flow exhibits a pronounced hysteresis behavior, so that different unsteady states can be observed for the same flow parameters. In this range multiple states may develop and co-exist, depending on the initial conditions.

Introduction. – Magnetohydrodynamic (MHD) 2 flows in ducts play a major role in liquid metal blankets for thermonuclear fusion reactors serving for cooling of the 3 reactor and for breeding tritium [1]. A very important re-4 quirement to these flows is that they should provide suffi-5 cient mixing of the fluid to improve heat and mass transfer 6 characteristics, i.e. provide sufficient level of turbulence. This, however, may be problematic as the liquid metal 8 flow occurs in a high magnetic field of about 5-10 Tesla. This leads to high MHD interaction between the induced 10 currents j and the magnetic field B resulting in a strong 11 braking Lorentz body force $j \times B$, which tends to damp 12 turbulence. 13

More often than not the ducts have thin electrically conducting walls, which leads to a velocity profile, shown in Fig. 1. Here the basic velocity exhibits a flat core, exponential Hartmann layers at the walls transverse to the magnetic field (the Hartmann walls), and jets at the walls parallel to the field (the side walls). Jets appear due to induced electric current within the core of the flow. The current must turn (either fully or partially) in the direction of the magnetic field at the side walls to complete the loops, which then close through the conducting Hartmann walls (Fig. 1). Once this happens, the resulting force $\mathbf{j} \times \mathbf{B}$ and its braking effect are reduced in the near-wall region, thus allowing the fluid to flow with higher velocity than in the core. The exact ratio of the core to jet velocities and thus between the mass fluxes strongly depends on the electrical conductivities of the Hartmann- and side-walls. In the extreme case of perfectly electrically insulating side walls and perfectly conducting Hartmann walls (known as Hunt's flow, see a particular case in [2]), practically all the mass flux is carried by the side-wall jets.

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We will be concerned here with an electrically conducting, incompressible fluid flow in a square duct, driven by a mean pressure gradient, with kinematic viscosity, ν , density, ρ , and electrical conductivity, σ . The flow is subjected to a uniform magnetic field of intensity *B* applied in the z-direction (x, y, z) are Cartesian coordinates, as shown in Fig. 1). For a given mean velocity *U*,



Fig. 1: Flow geometry of Hunt's flow in a square duct. Laminar velocity distribution with perfectly conducting Hartmannand perfectly insulating parallel-walls is shown at Ha = 100(blue); also shown are the streamlines of electric current density (brown).

the flow is characterised by three dimensionless parame-41 ters [3]: the Reynolds number, $Re = UL/\nu$, where L is 42 half the distance between the walls, the Hartmann num-43 ber $Ha = BL(\sigma/\rho\nu)^{1/2}$, the square of which characterises 44 the ratio between the Lorentz and viscous forces, and the 45 wall conductance ratio $c = \sigma_w h_w / \sigma L$, where σ_w and h_w 46 are the electrical conductivity and thickness of the wall, 47 respectively. Note that c may be different for each wall. 48

A family of the exact solutions for MHD flows with jets 49 was obtained in [2]. Early experimental results were re-50 viewed in [4], and the main conclusion was that the jets 51 were much thicker than the $O(Ha^{-1/2})$ thickness of the 52 jets in the fully developed flow. Linear stability of a flow 53 in a duct with c = 0.07 [5] could not account for this 54 increase in the thickness as the instability was fully con-55 tained in the parallel side layers. In addition, the Ting and 56 Walker (TW) vortices appeared for critical Reynolds num-57 ber $Re_{cr} \sim 300$, while in the corresponding experiment [6] 58 the instabilities appeared at $Re_{cr} \sim 3000$. For Hunt's flow 59 linear stability analysis was performed by [7], who discov-60 ered a rich variety of perturbations for Ha < 45, while 61 for higher Ha the most unstable mode consisted of TW 62 vortices. The controvercies between the theory and the 63 experiment were resolved in [8], where for Ha = 200 and 64 c = 0.5 direct numerical simulation (DNS) of the tran-65 sition to turbulence were performed. First, it was shown 66 that, when TW vortices appear, they are so weak that they 67 do not affect the mean velocity profile. For Re > 3500 a 68 new effect appeared consisting of partial jet detachment 69 from the side walls, which indeed led to thicker jets. 70

In this letter we will focus on the transition to turbulence in Hunt's flow and present various flow structures
obtained during the transition. Detailed analysis of these
state will be presented later in a series of journal papers.

Procedure and parameters. – We consider the MHD equations for incompressible fluid in a quasi-static form, which is standard for liquid metals (see, e.g., [3,5,8]). The governing equations are solved with second-order finite differences using our in-house solver [9]. The boundary conditions are periodic in the streamwise *x*-direction. Time integration is either fully explicit (using the 2^{nd} order Adams-Bashforth/Backward-differencing scheme) or with the implicit treatment of the viscous terms $\frac{1}{Re}\nabla^2 \boldsymbol{v}$. The semi-implicit version is beneficial at low-Re, as it helps to maintain large integration time-steps δt without the loss of numerical stability.

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The simulations have been conducted in a domain size of $8\pi \times 2 \times 2$ and numerical resolution varied from 512×128^2 to 1024×256^2 points in, correspondingly, *x*-, *y*- and *z*-directions. The computational grid is clustered in the wall-normal *y*- and *z*-directions by applying the coordinate transformation based on the hyperbolic tangent (see, e.g., [9]). Constant time step δt was maintained at $5 \times 10^{-4} \dots 2 \times 10^{-3}$ (depending on *Re*). The typical runtime contributed to $5 \times 10^5 \dots 10^6$ steps (several hundreds of convective units) using 128...1024 cores, depending on the resolution.

A series of DNS has been performed at a fixed Hartmann 97 number Ha = 100, representing the case of moderate mag-98 netic fields, and a broad range of Re, covering regimes from 99 Re = 200 to 10000. A similar approach with fixed Ha 100 was also used in [8]. In various studies on flow transition 101 the most commonly used scenario is to modulate the ini-102 tially laminar flow by perturbations. These can be either 103 random noise or a combination of specially constructed 104 modes modulated by random noise (one- or two-step sce-105 nario, e.g., [10]). In either case the transient evolution 106 and the actual thresholds are governed by many parame-107 ters, such as the amplitude of perturbations, specific spa-108 tial shape and distribution, optimal wave-numbers, etc. 109 To avoid this ambiguity, two other scenarios of obtain-110 ing time-dependent MHD solutions have been adopted. 111 First, we have used fully developed non-MHD turbulent 112 flows, precomputed for the closest possible Re, as the ini-113 tial state. Then the magnetic field is switched on and the 114 simulation is continued further. This method has the ap-115 peal that fluctuations are created in a natural way. In 116 another procedure the unsteady MHD solution, obtained 117 for a certain Re, is used as the initial state to study evo-118 lution to a different target Re, which can be either higher 119 or lower than the initial Re. This method has been most 120 extensively used to study the phenomenon of hysteresis. 121 We have employed both ways of changing Re – with small 122 steps and as a single shot to the target value. The onset 123 of new unsteady solutions or transition between different 124 states have been identified by visual analysis of the flow 125 field and by monitoring the kinetic energy of the trans-126 verse velocity components $q_t = u_y^2 + u_z^2$ (see, e.g., [11]). To attain a statistically sustained (fully developed) state 127 128 the simulations have typically run for the time from at 129 least 100 to 1000 convective time-units. 130

Results. –

Low- and mid-range values of Re. We begin our discussion with the results of simulations at Re = 500 shown in Fig. 2a. This flow state was obtained using turbulent non-MHD flow at Re = 1200 (the smallest Re where turbulence can still exist) as an initial condition. One interesting observation is that the turbulent energy drops by several orders of magnitude as the magnetic field is
switched on. Basically, the magnetic field rapidly destroys
turbulence so that the remaining perturbations can be
viewed as a residual noise. As a minimum is attained,
the residual fluctuations start to grow and finally evolve
into a different unsteady regime. This behavior has been
found in our simulations performed at other values of *Re*.

Here we can clearly observe inherently small vortical 145 structures located within the side layer region (Fig. 2a) 146 and insert). The vortices exhibit anisotropy in the vertical 147 z-direction, especially pronounced close to the Hartmann 148 walls, as illustrated by the isosurfaces of λ_2 -criterion. Ac-149 cording to [12], negative regions of λ_2 identify coherent 150 vortical structures (λ_2 is the 2^{nd} eigenvalue of the ten-151 sor $S_{ik}S_{kj} + \Omega_{ik}\Omega_{kj}$, where S_{ij} and Ω_{ij} are the rate of 152 strain and vorticity tensors). The structures are comprised 153 of clockwise rotating vortices (CW, thinner ones at the 154 side walls) and counter clockwise rotating vortices (CCW, 155 thicker ones closer to the core). Hereafter the direction of 156 vortex rotations is conventionally defined in respect to the 157 right-side if viewed along the flow direction (note the sign 158 change at the opposite wall). Both CW and CCW vortices 159 are organized in a slightly staggered arrangement. These 160 are the so-called TW vortices [5], they appear at low-Re161 and are viewed as the 1^{st} unstable solution of Hunt's flow. 162 The TW vortices are very weak, e.g., for this case their 163 kinetic energy is about 0.033% of the total kinetic energy 164 of the flow (Table 1), so hardly producing any visible im-165 pact on the integral flow charachteristics, such as the wall 166 stresses τ_w or the total friction coefficient C_f . 167

Further increase of Re reveals an interesting and rather 168 unexpected behavior of the flow. Namely, using this state 169 as the initial condition and increasing Re to 1000, we 170 have observed that TW vortices very quickly and nearly 171 completely disappear during the initial phase of simula-172 tion. However, after about a hundred of convective units 173 of temporal evolution, a new type of vortical structures 174 is developed. These new instabilities are much larger, 175 both in size and kinetic energy (about 0.1% of the to-176 tal kinetic energy), are elongated in the flow direction and 177 tend to dominate about 50...60% of the flow domain (Fig. 178 2b). This particular behavior is observed in the range 179 $1000 \leq Re \leq 1300$, where these big vortical structures 180 have also been repeatedly reproduced in simulations initi-181 ated by non-MHD turbulent states. 182

Increasing Re to 1400, we again observe another flow 183 regime. New structures develop in the form of additional 184 small-scale CW and CCW rotating vortices, which are lo-185 cated within the large instabilities previously observed at 186 $1000 \leq Re \leq 1300$, which evolve in this case too. There 187 are quite a few remarkable features concerning these struc-188 tures. On the initial inspection, they are very small in size 189 and elongated in the vertical z-direction, as demonstrated 190 in Fig. 2c and in Table 1. The shape resembles that of 191 "pike-teeth", which are settled in a staggered arrangement 192 in respect to the mid-plane symmetry z = 0. Further anal-193 ysis shows that the kinetic energy of these new structures 194

is about two times smaller than the previous instabilities 195 at Re < 1300, contributing 0.05% of the total kinetic en-196 ergy of the flow. However, the specific distribution of the 197 kinetic energy over velocity components changes. Albeit 198 the streamwise component is still dominating, the energy 199 associated with the vertical component increases by a fac-200 tor of 2, thus producing stronger anisotropy in the vertical 201 direction (e.g., q_z/q_u in Table 1). Basically, the "pike-202 teeth" can be viewed as rather strong (in amplitude) and 203 short (in the vertical length) modes. Once these structures 204 have developed, they remain in the flow throughout the 205 simulation. At higher Re (up to 1550) they develop more 206 rapidly, stretching in the vertical direction, and demon-207 strate an alternating behavior along the side walls. 208

To check the validity of the results we have conducted additional verifications. First, it has been found that these structures also arise if the simulation is initiated with turbulent state. Secondly, to make sure these structures were not a numerical artifact due to an insufficient grid, we have also performed simulations at higher resolution of 2048×384^2 points and found them to appear again. 205

The "pike-teeth" have been observed in a narrow range $1400 \le Re \le 1550$, beyond which they are followed by another type of instability – jet detachments, discussed in the next section. Given the narrow range of Re and the gradually increasing level of kinetic energy, it seems feasible that the "pike-teeth" may serve a nuclei of detachments. 221

Jet detachments and transition to turbulence. Upon 222 further increase of Re to 1630, we observe that the flow 223 exhibits another time-dependent regime - partial jet de-224 tachments. The temporal evolution and settlement of this 225 flow regime is shown in Fig. 3 with four phases, corre-226 sponding to the different times of flow evolution. One can 227 see that at the beginning a very small nuclei of instability 228 - a localized spot - appears at one side wall of the duct 229 (a). As time evolves, this spot grows in size and attains 230 kinetic energy from the mean flow until the side jet visibly 231 detaches from the wall (b). Soon after, the detached struc-232 ture interacts with the opposite side layer, thus, producing 233 a series of small disturbances (c), which rapidly evolve into 234 similar-sized patterns. Consequently, both walls become 235 populated with detached structures (Fig. 3d). 236

The nature of jet detachments can be understood more 237 thoroughly from the analysis of vortex patterns formed at 238 the side walls, as shown in Fig. 4 for Re = 2000...5000. 239 The plot at Re = 2000 (Fig. 4a) demonstrates small CW 240 rotating vortices, housed at the inner (near wall) region 241 of the domain of the high velocity jet. They are the re-242 sult of high shearing effects within the inner region of the 243 duct. Simultaneously, CCW vortices are formed between 244 the outer region of the velocity jet and the core flow. Ini-245 tially, at $Re \leq 1600$, the small CCW vortices remain stable 246 travelling in the streamwise direction along with the bulk 247 of the flow. At some point the CCW vortices develop both 248 in size and intensity and have a retarding effect on the jet 249 velocity in the outer region. Eventually the CCW vortices 250



Fig. 2: Instantaneous snapshots of flow states at Re = 500 (a), 1000 (b) and 1400 (c). Shown are isosurfaces of the second eigenvalue λ_2 of tensor $S_{ik}S_{kj} + \Omega_{ik}\Omega_{kj}$ (gold) and two isosurfaces of the spanwise velocity component u_y – positive (pink) and negative (cyan) of the same magnitude. The insert shows structure of CW and CCW vortices for Re = 500 in the mid-plane z = 0 using fluctuating velocity v, rotation is defined in respect to the right-side if viewed along the flow direction.

Re	regime	q/Q_{tot}	q_t/Q_{tot}	q_y/q	q_z/q	regime	q/Q_{tot}	q_t/Q_{tot}	q_y/q	q_z/q
	Incremental simulations					Decremental simulations				
300	laminar					jet detach.	2.8%	0.45%	13.7%	2.4%
500	TW vort.	0.033%	0.0081%	11.0%	13.5%	jet detach.	6.3%	1.2%	16.1%	2.9%
1000	large vort.	0.10%	0.0125%	4.0%	8.5%	jet detach.	9.5%	2.1%	18.0%	4.0%
1400	pike-teeth	0.05%	0.0110%	4.0%	18.0%	jet detach.	not performed			
2000	jet detach	. 10.0%	2.7%	19.0%	8.0%	jet detach.	same as incremental run			
5000	turb. jets	3.5%	1.5%	17.0%	26.0%	turb. jets	same as incremental run			

Table 1: TKE of velocity perturbations, shown in respect to the total kinetic energy of the flow Q_{tot} and as distribution over different components of the velocity fluctuations. Listed are the TKE of all perturbations $q = \langle u_x^2 + u_y^2 + u_z^2 \rangle$ and the transverse part $q_t = q_y + q_z$, where $q_y = \langle u_y^2 \rangle$ and $q_z = \langle u_z^2 \rangle$. The brackets $\langle \rangle$ stand for volume averaging, u_x is without mean flow, corresponding flow regimes are also indicated.



Fig. 3: Temporal evolution of jet detachments at Re = 1630and Ha = 100. Full streamwise velocity U_x is shown in the (x, y)-mid-plane at z = 0, corresponding to four phases of evolution (every 7 convective units), beginning from a single event on one side-wall to the fully developed unsteady behavior (from top to bottom). Flow direction is from left to right, magnetic field vector is perpendicular to the (x, y)-plane.

have developed sufficiently enough to dramatically alter 251 the flow regime and lift the jet away from the side wall 252 in the y-direction, thus, the first jet detachment occurs. 253 Once a detachment has occurred, its evolution pushes the 254 large CCW vortex through the core flow and hits the op-255 posite layer. Here it interacts with CCW vortices on the 256 opposite wall, thus generating one more detachment. The 257 phenomenon of jet detachments is not entirely new, it has 258 been observed in hydrodynamic flows [13]. 259

Further analysis has been conducted to address the dy-260 namics of jet detachments and to explore the evolution 261 towards even higher Re. We have performed a series of additional simulations and found that a staggered arrange-263 ment of partial jet detachments, as in [8], seems to be 264 the preferred pattern. At the same time, the number of 265 structures detached from the side walls is rather ambigu-266 ous: not only may it depend of the initial conditions, but 267 also may change during the simulation. In some cases, we 268 have observed non-symmetric flow states with only one wall populated by the spots, whereas the opposite layer 270 remained unaffected for an extended period, sometimes 271 for more than a hundred of convective time units. The 272 evolution at increasing Re (Fig. 4b, c) indicate that the 273 jet detachments gradually lose their characteristic form 274 and develop small-scale structures in their trailing tail. It 275 is clearly evident that already at Re = 4000 the side jets 276 have become almost turbulent, albeit rare detachments are 277 still present. Finally, at Re = 5000 (Fig. 4d) the side wall 278 regions become entirely turbulent. A remarkable feature 279 is the near stabilization of the core flow as Re increases, 280 which is particularly well observed at $Re \geq 5000$. 281

Multiple solutions. In the previous sections we have 282 demonstrated various flow regimes and, correspondingly, 283

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Fig. 4: Transition of states with jet detachmens towards fully tubulent side-wall jets shown for Re = 2000 (a), 3000 (b), 4000 (c) and 5000 (d). Instantaneous vortical patterns in the (x, y)-mid-plane are visualized by the streamlines of fluctuating velocity field v. Color gradient is highlighted by the magnitude of full-scale streamwise velocity component U_x . Flow direction is from left to right, magnetic field vector is perpendicular to the (x, y)-plane.

the appearance of different types of instabilities, as Re is 284 increased as a single shot or, alternatively, as the mag-285 netic field is instantly switched on. Here we study the 286 flow evolution with yet another approach - changing Re287 in steps, applying either increments (moving up) or decre-288 ments (moving down the Re axis). For the incremental 289 simulations, the initial state at Re = 500 and Ha = 100, 290 populated with weak TW vortices, have been used. Intrin-291 sically, during the incremental runs we have not found any 292 remarkable differences to the previously obtained results. 293 All flow regimes and the corresponding types of instabil-294 ities have been identified at essentially the same ranges 295 of Re, e.g., elongated vortices at Re = 1000...1100, "pike 296 teeth" at 1400, beginning of jet detachments at 1600...1700 297 and turbulization of the side jets at Re > 4000. 298

The situation changes in reverse simulations, starting 299 from a state of Re = 10000 with turbulent side jets. The 300 initial and relatively long part between 10000 $\geq Re \geq$ 301 2000 demonstrates the reverse sequence of flow regimes 302 vs. the incremental runs, i.e. gradual transition from tur-303 bulence to jet detachments (Table 1). However, a contin-304 ued further decrease of Re reveals no appearance of other 305 intability types, identified in the incremental simulations. 306 Instead, unstable flow regimes with jet detachments con-307 tinue until $Re \sim 200$, albeit the kinetic energy of pertur-308 bations gradually decreases. This behavior demonstrates 309 the phenomenon of multiplicity of possible states and so-310 lutions, which appear depending of the particular route. 311 A remarkable feature is that this behavior has been ob-312 served in a very broad range 200 < Re < 2000, inherently 313 in the parameter space varying by one order of magni-314 tude. Similar hysteresis effects were also observed in the 315 prior study of duct flow with finite wall conductivity [8]. 316 However, the range of Re with multiple states was far 317 more narrow. An example of such co-existing solutions is 318 shown in Fig. 5 for two different flow states at the same 319 Re = 1000. We can see both the pattern typical for CW 320 and CCW vortices (Fig. 5, top) and jet detachments (Fig. 321 5, bottom) obtained in, correspondingly, incremental and 322 decremental simulations. Our further investigations have 323 shown that the multiplicity is not only affected by the di-324

Fig. 5: Instantaneous snapshots of flow fields at Re = 1000 shown for incremental (top) and decremental (bottom) simulations. The flow patterns are visualized by the vertical vorticity component ω_z in the (x, y)-mid-plane ("blue-red" gradient corresponds to "negative-positive" ranges of ω_z).

rection of varying Re, but also by the initial conditions (e.g., uncorrelated turbulent non-MHD states).

This figure also demonstrates another interesting effect – the phenomenon of vortex shedding in the form similar to that of a Kármán street. We have found that this effect is predominantly expressed at low-Re range. The motion of detached structures through the domain generates a vortex shedding in their downwind, thus, producing patterns similar to flows past a solid obstacle.

Summary. – The instability of Hunt's flow has been 334 studied numerically at a fixed Ha = 100, which corre-335 sponds to moderate magnetic fields, and a broad range of 336 Re varying from 200 to 10000. Upon increasing Re sev-337 eral unsteady flow regimes have been identified, includ-338 ing the TW vortices (low Re < 1000), partial jet detach-339 ments (mid-range of $Re \geq 1630$) and transition to fully 340 developed turbulence in the side-wall jets (upper range 341 of $Re \geq 5000$). In addition, two new instabilies are dis-342 covered: large elongated vortices (1000 < Re < 1400)343 and very small, tightly localized at the side walls, vertical 344 "pike-teeth" structures (1400 < Re < 1550). The results 345 of our simulations suggest that these structures can be 346 viewed as transients, connecting the other, major types 347 of unstable solutions. Indeed, these two new regimes are 348 found in rather narrow ranges of Re, being quickly fol-349 lowed by, or proceeded by another persistent unsteady so-350 lution, when *Re* slightly changes. 351

Our simulations also suggest another view on the in-352 stability of Hunt's flow. Namely, given the extremly low 353 amplitude of TW vortices observed at low-Re, one may 354 speculate that the actual transition to the truly unsteady 355 flow states should be viewed beginning from the appear-356 ance of jet detachments. By adopting this point of view, 357 we can see the laminar-turbulent transition in Hunt's flow 358 is not a unique feature, but is rather similar to other MHD 359 and non-MHD shear flows. Indeed, approaching the crit-360 ical range of $Re \approx 1600$, detachments first appear in the 361 form of sporadic events, i.e. essentially isolated spots re-362 siding in the side wall layers. As *Re* increases, the side lay-363 ers become increasingly populated with such spots, which 364 keep developing small-scale fluctuations, until the entire 365 extent of the side layer is involved into fully turbulent mo-366 tion, as demonstrated in Fig. 6. Very similar flow dynam-367 ics with puffs bordering laminar and turbulent regimes is 368 well known for many other shear flows, in particular for 369



Fig. 6: Evolution of jet detachments versus Re, instantaneous snapshots of flow fields are visualized at Re = 2000 (a), 3000 (b) and 4000 (c). Shown are the isosurfaces of full-scale streamwise velocity U_x at the side walls (light blue), TKE of transverse velocity $u_y^2 + u_z^2$ (gold) and λ_2 criterion (cyan). The inserts show velocity profiles in the (y, z)-section at $x = L_x/2$.

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MHD duct and pipe flows with insulating walls [11], where 370 the puffs are localized at the side walls. Even the non-371 symmetry of pattern arrangements at the opposite side 372 walls is observed too. The clearly novel feature of Hunt's 373 flow instability is that these transient states are observed 374 in a much broader range of *Re*, distinguishing it from other 375 shear flows. The same observation also applies to multiple 376 states and hysteresis, found in an extremely broad range 377 2000 > Re > 200.378

At the early stages of jet detachments, the vortical structure detaching from the wall can also produce the effect very similar to that of an obstacle, which is observed since the velocity of the core flow is much smaller than in the jet. As a result, the phenomenon of vortex shedding may form.

Another distinct feature we have observed is the non-385 monotonic behavior of the core flow versus increasing 386 Re number: essentially unperturbed core flow at low-Re 38 regimes (TW vortices), evolution into a strong unsteady 388 motion at moderate Re (jet detachments) and approaching 389 almost unperturbed state, populated with weak quasi-2D390 structures at high Re (turbulent side-wall jets). This par-391 ticular behavior of the core flow has practical implications. 392 Namely, the most unsteady regimes at intermediate values 393 of Re seem the most preferred ones to intensify flow mix-394 ing and, correspondingly, enhance heat transfer, which is 395 the ultimate purpose of fusion blankets. 396

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